

Hints for leptonic CP violation or New Physics?

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(Dated: January 18, 2016)

One of the major open questions in the neutrino sector is the issue of leptonic CP violation. Current global oscillation data shows a mild preference for a large, potentially maximal value for the Dirac CP phase in the neutrino mixing matrix. In this letter, we point out that New Physics in the form of neutral-current like non-standard interactions with real couplings would likely yield a similar conclusion even if CP in the neutrino sector were conserved. Therefore, the claim for a discovery of leptonic CP violation will require a robust ability to test New Physics scenarios.

In 2015 a Nobel prize for the discovery of neutrino oscillation was given. The bulk of the existing data currently is very well described by the oscillation of three active neutrinos, see for instance Ref. [1, 2]. There are potential indications for additional, sterile neutrinos, which at this stage are not conclusive yet [3] and we therefore will neglect for the remainder of this letter. Lately, due to a tension between reactor and accelerator neutrino experiments, a preference for a value of the Dirac CP phase close to $-\pi/2$ was first reported in Ref. [4]. The question of CP violation in the leptonic sector is a priority of the future neutrino program and the main effort is dedicated to the DUNE experiment [5]. Apart from a number of experimental challenges and the need to understand neutrino-nucleus interactions, sub-leading effects of theoretical origin can also affect the determination of the Dirac CP phase. In this letter we focus on so-called non-standard interactions (NSI), which have been speculated about even before the discovery of neutrino oscillation [6] and for a recent review see Ref. [7]. NSI provide a model independent, effective field theory framework to include new physics effects in the standard neutrino description, see *e.g.* Ref. [8].

NSI are parameterized by dimensionless couplings, ε , which are measured relative to G_F . From neutrino data alone large magnitudes for the dimensionless couplings $|\varepsilon_{e\tau}^{f=u,d}| \lesssim 0.14$ at 90% [9] are allowed, which for Earth matter densities translates into $|\varepsilon_{e\tau}| \sim \mathcal{O}(1)$. Generally, NSI can provide new phases that can potentially be new sources of CP violation. In the context of flavor changing NSI in the source and the detection of neutrinos a discussion of new sources of CP violation appeared in Ref. [10]. In the case of neutral current NSI a discussion of this issue appeared in Ref. [11]. In the context of NOvA, and motivated by large NSI allowed by solar neutrino data, one example of the potential NSI and standard oscillation confusion was analyzed at the probability level in Ref. [12].

Direct bounds derived from the neutrino sector are typically of the order 0.1-1 in units of G_F , that is, New Physics contributions of about the same size as the leading Standard Model (SM) contribution are still allowed. On the other hand, any model where these new inter-

action are introduced above the electroweak symmetry breaking scale has to address the fact that invariance under the weak $SU(2)$ group will create a charged lepton counter part of any neutrino-only operator. Given that the electroweak scale is not very far from the scale at which typical neutrino experiments are conducted, breaking of the electroweak symmetry does not erase the correspondence between neutral and charged lepton operators, at best factors of a few are gained. The charged lepton bounds involving the first and second family are very stringent. In Ref. [13] a systematic analysis of dimension 6 and 8 operators is performed and it is found that for dimension 8 operators it is possible to arrange for cancellations, provided a suitable particle content is chosen, such that large NSI in the neutrino sector can be realized without creating sizable effects in the charged lepton sector. This requires fine tuning and no actual models have been put forward. It is worth noting that even without fine tuning, third family NSI in the τ sector can be potentially large of order $0.1 G_F$, see *e.g.* [14], since the corresponding charged lepton bounds themselves are very weak.

The situation is quite different if we consider models where NSI is generated *below* the scale of electroweak symmetry breaking, since by construction there will be no correspondence between neutral and charged lepton operators. An early example of a low-scale neutrino mass model (without NSI) is given in Ref. [15], where the breaking of a discrete gauge symmetry at a scale as low as a few keV is responsible for small Dirac neutrino masses. It would be straightforward to augment this model by additional flavor changing neutral currents to create large NSI. The general idea is to invoke new light degrees of freedom which preferentially couple to neutrinos and/or dark matter particles, *e.g.* [16, 17]. These models tend to introduce new self-interactions in the dark matter sector and need to observe the relevant astrophysical bounds from structure formation. Also, in some of these models there are connections to the sterile neutrino sector, which in turn can help to accommodate sterile neutrinos in cosmology [18]. Thus, we see that there is ample room from both an experimental and theoretical perspective for relatively large NSI. The resulting degeneracies between New

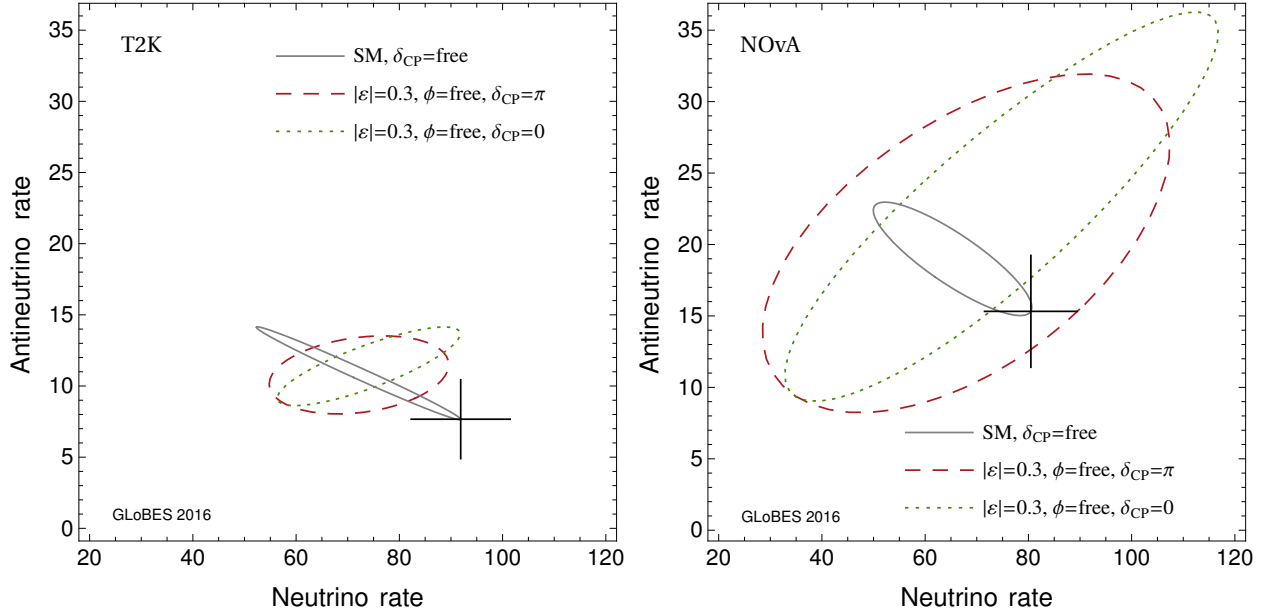


FIG. 1. Bi-rate plots. The full line curve corresponds to the SM for all Dirac CP phase values. The dashed and dotted curves correspond to a fixed NSI magnitude $|\varepsilon| = 0.3$ and for all NSI phase ϕ values but for different values of the Standard CP phase. The cross for the SM point $\delta_{CP} = -\pi/2$ shows the statistical uncertainty. The left (right) panel corresponds to our implementation of the T2K (NOvA) experiment. All parameters not labeled in the plot were fixed to their best fit values, see text for details.

Physics and oscillation physics recently have been studied in Ref. [19] in a more general setting. Here, we will study specifically the impact neutral current-like NSI can have on the analysis of Daya Bay, T2K and NOvA data sets and point out that the current hint for maximal CP violation maybe in fact be caused by CP *conserving* New Physics.

The standard neutrino oscillations in vacuum are described by the Hamiltonian:

$$H_0 = \frac{1}{2E} [U \text{diag}(0, \Delta m_{21}^2, \Delta m_{31}^2) U^\dagger] \quad (1)$$

where U is the lepton mixing matrix parameterized by three mixing angles θ_{ij} and a CP violating phase δ_{CP} . Δm_{k1}^2 in Eq. (1) denotes the two known mass square differences and E the neutrino energy.

Since we will consider long-baseline neutrino oscillations, the neutrino forward scattering interactions in matter, can be effectively parameterized in the presence of NSI by the following Hamiltonian:

$$H_{\text{int}} = V \begin{pmatrix} 1 + \varepsilon_{ee} & \varepsilon_{e\mu} & \varepsilon_{e\tau} \\ \varepsilon_{e\mu}^* & \varepsilon_{\mu\mu} & \varepsilon_{\mu\tau} \\ \varepsilon_{e\tau}^* & \varepsilon_{\mu\tau}^* & \varepsilon_{\tau\tau} \end{pmatrix} \quad (2)$$

with $V = \sqrt{2}G_F N_e$, where G_F is the Fermi constant and N_e is the electron density on Earth. Notice that the Hamiltonian in Eq. (2) has 8 new physical parameters in addition to the standard oscillation ones. However, only few of them are present in an specific oscillation channel.

In the particular case of the (anti)neutrino appearance channel, only the two NSI complex parameters $\varepsilon_{e\mu}$ and $\varepsilon_{e\tau}$ play a role [20]. Instead of an exhaustive analysis to quantify the interplay of all the NSI parameters with the SM ones, we will consider $\varepsilon_{e\mu} = 0$ and $\varepsilon_{e\tau} \equiv |\varepsilon| \exp(i\phi)$ which is enough for our discussion. Throughout this letter, our results will correspond only to the normal ordering for the neutrino spectrum.

We have implemented a GLOBES [21, 22] simulation of a 295 km baseline neutrino beam experiment with the characteristics of T2K but scaling its exposure by a factor 5 relative to the current data [23]. In addition, we have also implemented a simulation of NOvA running 3 years in neutrino mode plus 3 years in antineutrino mode [24]. The set of oscillation parameters used along this work corresponds to the best fit values in Ref. [1] except for the reactor mixing angle that was fixed to the Daya Bay best fit value in Ref. [25].

The interplay of the Dirac CP phase and the NSI parameters in each of the two considered facilities is shown in the bi-rate plots of Fig. 1. For the chosen values of the NSI parameters, and considering the errors in both NSI and SM rates, an overlap in the SM point $\delta_{CP} = -\pi/2$ is clearly shown in both panels for T2K and NOvA. However, notice that in the case of the T2K rates (left panel) there is a tension between the considered NSI cases and the the SM point $\delta_{CP} = -\pi/2$. Thus, in combination with NOvA the ‘confusion’ will be significant for the NSI case with $\delta_{CP} = \pi$.

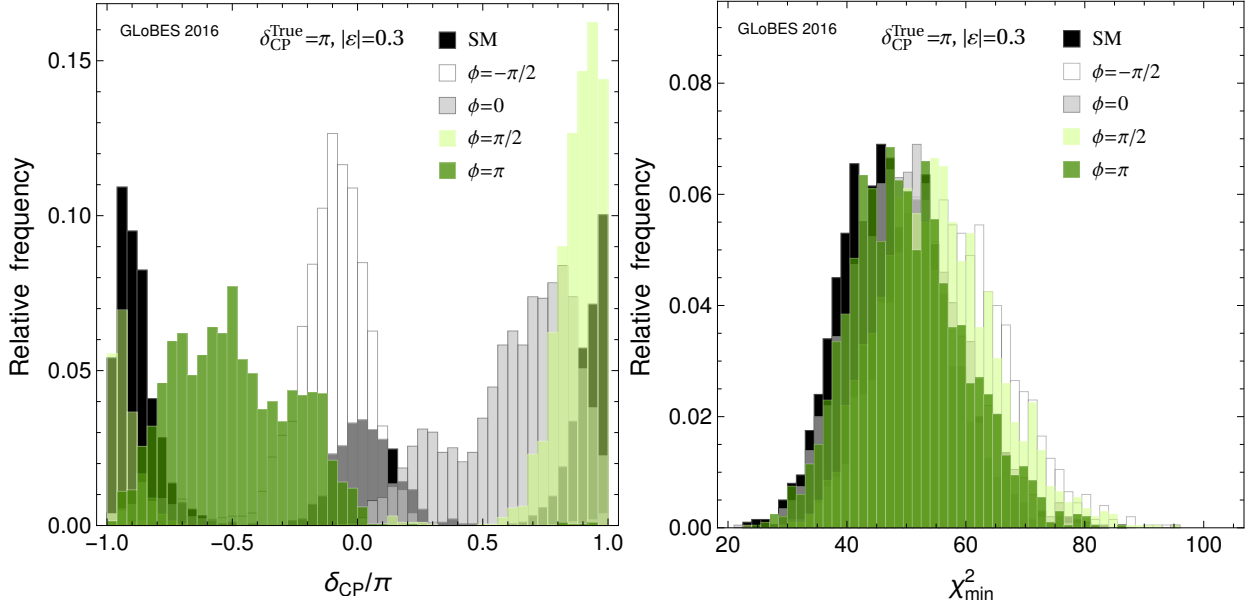


FIG. 2. Results assuming $\delta_{CP}^{True} = \pi$. In the left panel, the Dirac CP phase best fit distributions for Standard Model (SM) and NSI interactions are shown. The NSI magnitude was fixed to the value $|\epsilon| = 0.3$ for the different NSI phase ϕ values showed in the plot. In the right panel appears the minimum χ^2 distributions from the fit to the SM Dirac CP phase for each of the SM and NSI cases showed in the left panel. All not shown parameters were marginalized over, see text for details of the analysis.

By assuming CP conserving values, in the presence of NSI, the freedom in $|\epsilon|$ and ϕ can be used to ‘mimic’ the SM point $\delta_{CP} = -\pi/2$. Basically, the value of $|\epsilon|$ sets the opening of the NSI ellipse while the phase ϕ can be tuned to coincide with the SM and NSI intersection point. Notice that, for T2K (left panel), a larger value for $|\epsilon|$ is needed to exactly pick the SM point $\delta_{CP} = -\pi/2$ while for NOvA (right panel) a smaller ϵ would instead be required. This is mainly due to the different baseline of both experiments since the sensitivity to the NSI depends on the SM and NSI interference of vacuum and matter oscillations (see Eq. (33) of Ref. [20]). Thus, a combination of experiments with different baselines limits the parameter tuning we are discussing, which in the future may allow to disentangle these effects if both high-statistic data sets from DUNE and T2HK are available [26–28]. However, our philosophy here is quite different: Given a particular set of NSI parameters (allowed by current bounds) we want to quantify the level of ‘confusion’ if the currently preferred value for the δ_{CP} turned out to be correct.

To quantify the sensitivity to the SM Dirac CP phase in the presence of NSI we have adopted the usual χ^2 analysis where $n_i = n_i(\vec{\lambda}^{true}, \epsilon^{true})$ and $\mu_i = \mu_i(\vec{\lambda})$ corresponds to the simulated and test events, respectively. The set of standard oscillation parameters, mixing angles, Dirac CP phase and the solar and atmospheric splittings, are represented by $\vec{\lambda}$. To test SM oscillations rates μ_i against the ones with NSI, we implemented random statistical fluctuations (Poisson distributed) on the ‘true’ rates n_i . Gaussian priors for all the standard oscillation param-

eters but δ_{CP} were included in the analysis.

Our main result is presented in Fig. 2. The distribution of the Dirac CP phase best fit values for SM and NSI interactions are shown in the left panel. In all cases $\delta_{CP}^{true} = \pi$. In the case of the SM, the δ_{CP} has values distributed around CP conserving values including zero but it is mainly distributed around $\delta_{CP} = \pm\pi$ as expected. When NSI are considered, two general features appear: In the case of complex NSI, $\phi = -\pi/2$ and $\phi = \pi/2$, δ_{CP} is distributed with a higher probabilities respect to the SM case around values that are close to zero and π , respectively; and the distributions are very peaked around their mean value. However, when $\phi = 0$ (real NSI) the distribution of δ_{CP} is broader and extends for half of the parameter space in the positive region. In this case, the mean value is located near $\delta_{CP} = \pi$ with a probability closer to the one of the SM case. In the remaining case of real NSI $\phi = \pi$ also the distribution is broader but the main feature is that its mean value is centered at $\delta_{CP} = -\pi/2$ almost with the same probability of the SM case. In both cases with real NSI there is a non-negligible probability to find δ_{CP} violating values even though δ_{CP}^{True} has assumed to be CP conserving. This result is remarkable, in particular when $\phi = \pi$, in the light of the current preference for the Dirac CP phase value.

The corresponding distribution of the χ^2 -minima, for each case of the right hand panel, are shown in the left hand panel of Fig. 2. Due to the random statistical fluctuations the χ_{min}^2 for the SM case is centered at $\chi^2 \simeq 40$ corresponding the number of bins minus the number of fitted parameters. However, the main feature is that all

the NSI χ^2 distributions are centered within ~ 7 units from the SM central value with almost the same probability. In the main case of $\phi = \pi$ both SM and NSI the χ^2 distributions are even closer, they almost completely overlap each other. This result shows that existing experiment can *not* distinguish these two cases.

To summarize, we have studied the robustness of the recent hint for maximal leptonic CP violation in the presence of neutral current-like NSI. We simulate many iterations of the same experiments, T2K and NOvA, and perform a fit to the resulting data, both for a purely three flavor oscillation and in presence of NSI. We find, that even if CP is fully conserved by the NSI and standard oscillations, a preference for the best fit value of the leptonic CP phase $\delta \simeq -\pi/2$. Thus, the current hint for maximal leptonic CP violation can be either due to maximal leptonic CP violation or CP conservation in the presence of New Physics. The χ^2 distributions in both cases are nearly identical, high-lighting the need for new experiments like DUNE and T2HK to resolve this confusion.

We would like to thank P. Coloma for her help with the GLOBES files and J. Kopp for a clarification in the use of his NSI tool. This work was supported by the U.S. Department of Energy under award DE-SC0013632.

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